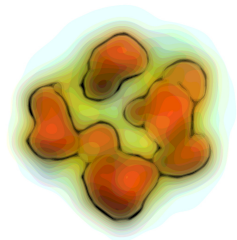


Inhomogeneity growth in two-component fermionic systems

P. Napolitani, M. Colonna

- 1) Can we apply **Fermi-liquid** theory of nuclear matter (**NM**) to **heavy-ion** collisions (**HIC**) in a **one-body** framework ?
- 2) Is the 3D solution of the BL eq. in NM giving **fluctuations of correct amplitude** in both **isoscalar/isovector** channels ?
- 3) New understanding of HIC ?

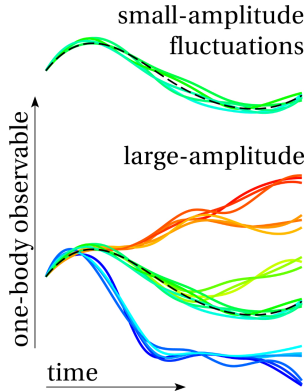


the full story :

[[NAPOLITANI, COLONNA ARXIV :1705.08268 2017](#)]

Describing large-amplitude dynamics of nucleonic systems

- *small amplitude-limit at low energy*
→ correlated channels, coherent states
[TDGCM : REINHARD,GOEKE REPPrPH50(1987), GOUTTE *et al* PRC 71 (2005),
BALIAN-VÉNÉRONI (1981,1992), SIMENEL EPJA 48 (2012),...]
- *large-amplitude regimes at low energy* →
→ non-correlated states beyond
single-part. picture [LACROIX ARXIV :1504.01499 (2015)]
- *excitation beyond Pauli blocking* →
→ observables width spread, dissipation
[ETDHF : WONG,TANG PRL 40 (1978), LACROIX *et al* PROGPartNucPhys 52 (2004),...]



- *chaotic regime, bifurcation*
→ highly non-linear, possibly unstable
[STDHF / BOLTZMANN-LANGEVIN : REINHARD,SURAU ANNPHYS 216(1992), SLAMA,REINHARD,SURAU ANNPHYS 355 (2015),...]
⇒ (1) Clusterisation from one-body density fluctuations
⇒ (2) n, p transport between fragments and the medium

Handling many-body correlations in mean-field extensions

- Pure mean-field equations are not valid in regions where instabilities, bifurcations, chaos are present

BBGKY (Born Bogoliubov Green Kirkwood Yvon) hierarchy

for a 2-body interaction V_{ij} :

$$i\hbar \frac{\partial \rho_1}{\partial t} = [k_1, \rho_1] + \text{Tr}_2[V_{12}, \rho_{12}]$$

$$i\hbar \frac{\partial \rho_{12}}{\partial t} = [k_1 + k_2 + V_{12}, \rho_{12}] + \text{Tr}_3[V_{13} + V_{23}, \rho_{123}]$$

kinetic equations

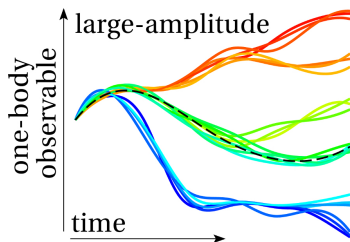
$$i\hbar \frac{\partial \rho_{1\dots k}}{\partial t} = \sum_{i=1}^k [k_i + \sum_{j<i} V_{ij}, \rho_{1\dots k}] + \sum_{i=1}^{k+1} \text{Tr}_{k+1}[V_{ik+1}, \rho_{1\dots k+1}]$$

high coupling regimes

($\rho_{1\dots k}$ =many-body dens. matrix, k_i =Kin.E operator, Tr_{ij} =partial trace)

- aim : going from 2nd order approx. to highly non-linear regimes
- higher-orders approximately recovered (rather than truncated) by handling **several mean fields**

[LACROIX *et al* EPJA52 (2016)]



Beyond 2nd order through a stochastic treatment

define a MF-trajectory subensemble $\{\rho_1^{(n)}; n = 1 \dots \text{sub}\}$ over $\tau_{\text{BL}} = t - t_0$ with small fluctuations around the mean

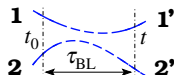
\Leftrightarrow probability $\rho_1^{(n)}, \rho_2^{(n)}$ to find two nucleons 1, 2 at two config. points not all time decorrelated

$\rho_{12}^{(n)}$ recovers some correlations of upper BBGKY orders, so that :

$$\rho_{12}^{(n)}(t_0) = \overbrace{\tilde{\Omega}_{12} \mathcal{A}_{12}(\rho_1^{(n)}(t_0) \rho_2^{(n)}(t_0)) \tilde{\Omega}_{12}^+}^{\text{collisional correlations}} + \overbrace{\delta \rho_{12}^{(n)}(t_0)}^{\text{fluctuations}} ; \quad \begin{array}{l} \text{introduces fluct.} \\ \text{around the} \\ \text{collision integral} \end{array}$$

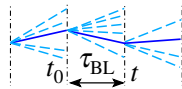
$$\langle \delta \rho_{12}^{(n)}(t_0) \rangle_{\tau_{\text{BL}}} = 0 ;$$

$$\langle \delta \rho_{12}^{(n)}(t_0) \delta \rho_{12}^{(n)}(t') \rangle_{\tau_{\text{BL}}} = \text{gain} + \text{loss}$$



(diffusion matrix : $G_{12} = V_{12} \tilde{\Omega}_{12} \rightarrow$ NN diff. cross section $|G_{12}|^2 \sim d\sigma/d\Omega$)

- if $\delta \rho_{12}^{(n)}(t_0) = 0 \Rightarrow \rho_1^{(n)}, \rho_2^{(n)}$ fully decorrelated \rightarrow 2nd order truncation without fluctuations \rightarrow (quantum) Boltzmann
- else $\Rightarrow \delta \rho_{12}^{(n)}(t_0)$ is an intermittent source of fluctuation seeds \rightarrow it can be exploited as a stochastic source



Obtaining an exploitable description

For one mean-field trajectory n in τ_{BL} :

Stochastic-TDHF scheme

$$i\hbar \frac{\partial \rho_1^{(n)}}{\partial t} \approx [k_1^{(n)} + V_1^{(n)}, \rho_1^{(n)}] + \overbrace{\bar{I}_{\text{coll}}^{(n)}}^{\text{average coll. term}} + \underbrace{\delta I_{\text{coll}}^{(n)}}_{\text{fluctuating coll. term}}$$

after τ_{BL} , it yields $\rho_1^{(n)} \rightarrow \{\rho_1^{(n,\lambda)}; \lambda = 1, \dots, \text{sub}_\lambda\}$

[REINHARD, SURAUD ANNPHYS 216 (1992); ANNPHYS 355 (2015)]

LACOMBE, REINHARD, SURAUD, DINH ANNPHYS 373 (2016)]

Boltzmann-Langevin One Body

$$\frac{\partial f^{(n)}}{\partial t} - \{h^{(n)}, f^{(n)}\} = I_{UU}^{(n)} + \delta I_{UU}^{(n)} = g \int \frac{d\mathbf{p}_b}{h^3} \int W_{(AB \leftrightarrow CD)} F_{(AB \rightarrow CD)} d\Omega$$

transition rate

occupancy

$$W_{(AB \leftrightarrow CD)} = |v_A - v_B| \frac{d\sigma}{d\Omega}; \quad F_{(AB \rightarrow CD)} = [(1-f_A)(1-f_B)f_C f_D - f_A f_B(1-f_C)(1-f_D)]$$

A, B, C, D : extended equal-isospin phase-space portions of size=nucleon imposed by the variance $f(1-f)$ in h^3 cells at equilibrium

[NAPOLITANI, COLONNA PLB726 2013; ARXIV :1705.08268 2017]

Boltzmann-Langevin

$f^{(n)}$: distribution functions

→ Fermi stat. at equilibrium

$$\frac{\partial f^{(n)}}{\partial t} = \{h^{(n)}, f^{(n)}\} + I_{UU}^{(n)} + \delta I_{UU}^{(n)}$$

Markovian contrib. :

$$\langle \delta I_{UU}^{(n)}(\mathbf{r}, \mathbf{p}, t) \delta I_{UU}^{(n)}(\mathbf{r}', \mathbf{p}', t') \rangle = \text{gain} + \text{loss} = 2\mathcal{D}(\mathbf{r}, \mathbf{p}; \mathbf{r}', \mathbf{p}', t') \delta(t-t')$$

[AYIK, GRÉGOIRE PLB212(1988); NPA513(1990)]

COLONNA, CHOMAZ, RANDRUP NPA567(1994)]

Wigner ↓ tr.



The same scheme for NM and open systems

*simplified SKM** [GUARNERA, COLONNA, CHOMAZ PLB373 (1996)]

$$\frac{E_{\text{pot}}}{A}(\rho) = \frac{A}{2}u + \frac{B}{\sigma+1}u^\sigma + \frac{C_{\text{surf}}}{2\rho}(\nabla\rho)^2 + \frac{1}{2}C_{\text{sym}}(\rho)u\beta^2$$

$A = -356 \text{ MeV}$, $B = 303 \text{ MeV}$, $\sigma = 7/6 \rightarrow K = 200 \text{ MeV}$ (soft);

$C_{\text{sym}}(\rho) = 32(\text{asy-stiff}) / \rho_{\text{sat}}(482 - 1638\rho) \text{ MeV}$ (asy-soft);

• momentum dependence omitted

residual term

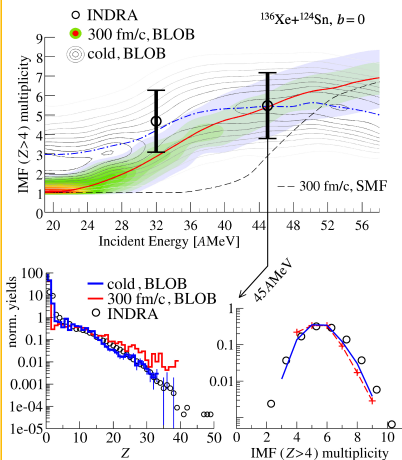
• E -dependent free σ_{NN} or screened BLOB) $\delta I \rightarrow$ fluct. in full phase space

we may compare to :

SMF) $\delta I \rightarrow$ separately treated as a stochastic force related to $U_{\text{ext}} \Rightarrow$ fluctuations projected on spacial ρ

HIC \Rightarrow fragment observables consistent with exp. data in BLOB \rightarrow

an example :



[COLONNA, NAPOLITANI, BARAN Nuc. CLUSTER CORRELATIONS... (2017),

DATA FROM INDRA : PRC86 (2012), EPJA 50 (2014)]

The same scheme for NM and open systems

*simplified SKM** [GUARNERA, COLONNA, CHOMAZ PLB373 (1996)]

$$\frac{E_{\text{pot}}}{A}(\rho) = \frac{A}{2}u + \frac{B}{\sigma+1}u^\sigma + \frac{C_{\text{surf}}}{2\rho}(\nabla\rho)^2 + \frac{1}{2}C_{\text{sym}}(\rho)u\beta^2$$

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• momentum dependence omitted

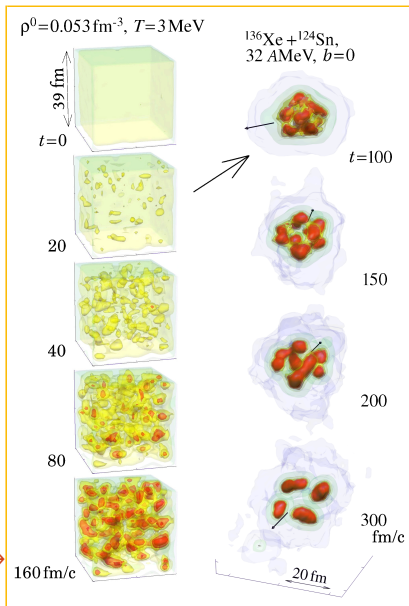
residual term

• E -dependent free σ_{NN} or screened BLOB) $\delta I \rightarrow$ fluct. in full phase space
SMF) $\delta I \rightarrow$ fluctuations projected

HIC \Rightarrow fragment observables consistent with exp. data in BLOB

NM \Rightarrow are fluctuation amplitudes also consistent with analytic expectations?

\rightarrow check in initially homogeneous NM \rightarrow (Fermi-Dirac at $\rho^0, T = 3$ MeV)



Fluctuations in nuclear matter and the BL equation

- Small disturbance in **uniform matter** around the mean trajectory f^0 :

$$\begin{aligned}
 q=s) \text{ isoscalar, } n, p \text{ in phase : } & \delta f^s = (f_n - f_n^0) + (f_p - f_p^0) \\
 q=v) \text{ isovector, } n, p \text{ out of phase : } & \delta f^v = (f_n - f_n^0) - (f_p - f_p^0)
 \end{aligned}
 \quad \delta f(\mathbf{r}, \mathbf{p}, t) \ll f^0(\mathbf{p}, t)$$

\Rightarrow fluctuation ρ_k^q of variance $(\sigma_k^q)^2$ at equil.

- BL eq. for a fluctuating field U' , in **symmetric uniform matter**, at low T ($\Rightarrow I_{UU} = 0$), applied to the disturbance δf^q (1st order approx) :

stable modes : Fluctuation-dissipation th.

spacial density correlation variance in ΔV at equilibrium (T)

$$(\sigma_k^q)^2 = \frac{T}{F^q(k)} ; \quad (\sigma_{\rho^q})^2 = \frac{T}{\Delta V} \left\langle \frac{1}{F^q(k)} \right\rangle_k$$

related to free-energy dens. curvature

- \Rightarrow *nucl. matter* : check $(\sigma_{\rho^v})^2$ versus Symmetry E
- \Rightarrow *open-system analogy* : isotopic distributions

unstable modes

$$(\sigma_k^q)^2(t) \approx D_k \tau_k (e^{2\frac{t}{\tau_k}} - 1) + (\sigma_k^q)^2(t=0) e^{2\frac{t}{\tau_k}}$$

both continuous and initial fluctuation seeds yield an exponential growth τ_k

[COLONNA et al PRC47(1993) NPA567 (1994)]

- \Rightarrow *nucl. matter* : check instability growth rates
- \Rightarrow *open-system analogy* : clusterisation

Isvector fluctuations in two-component nuclear matter

Isvector behaviour ($q \rightarrow v$, $\rho^v = \rho_n - \rho_p$) in BL in stable uniform matter

- scalar terms suppressed in $U^v \rightarrow$ stable conditions at all ρ_0

fluctuation-dissipation th.

$$(\sigma_{\rho^v})^2 = \frac{T}{\Delta V} \left\langle \frac{1}{F^v(k)} \right\rangle_{\mathbf{k}} \longrightarrow F_{\text{eff}}^v = \frac{T}{2\Delta V} \frac{\rho^0}{(\sigma_{\rho^v})^2} = \frac{T}{2\Delta V} \frac{\rho^0}{\langle [\delta\rho_n(\mathbf{r}) - \delta\rho_p(\mathbf{r})]^2 \rangle} \propto E_{\text{sym}}$$

[COLONNA PRL110 (2013)]

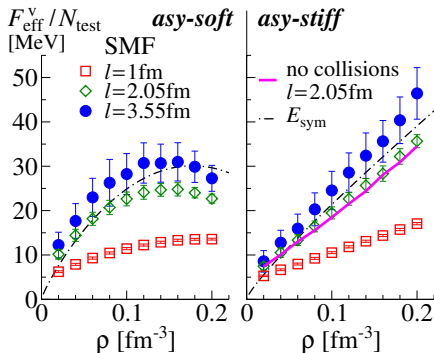
SMF ($\delta I \rightarrow$ projection) :

- solved with N_{test} test particles and fluctuation from MF noise

$$\Rightarrow F_{\text{eff}}^v \approx N_{\text{test}} E_{\text{sym}}$$

- same result without residual term
- \Rightarrow explicit iv term are missing in the iv channel

[NAPOLITANI, COLONNA ARXIV :1705.08268 (2017)]



l : edge of periodic box (surface effects if too small)

Isvector fluctuations in two-component nuclear matter

BLOB ($\delta I \rightarrow$ full ph. space) :

- larger iv variance, as a function of ρ^0
- small $\rho^0 \rightarrow$ ineffective coll. correlations
- large $\rho^0 \rightarrow$ longer path to convergence

[NAPOLITANI, COLONNA ARXIV :1705.08268 (2017)]

- noise issue : smearing from approx. MF mapping [REINHARD, SURAUD ANNPHY216 (1992)]

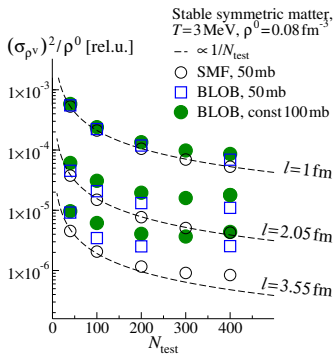
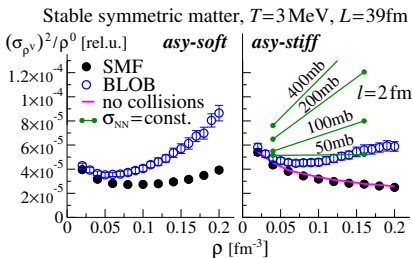
$\rightarrow N_{\text{test}}$ dependence. \rightarrow cured by :

(1) *reviving NN correlations*

(larger coll. rate)

(2) *augmenting N_{test}*

\Rightarrow still not enough in equilibrated NM (low T , low $\rho \rightarrow$ rare collisions), but better cured in early stages of HIC (out of equilibrium, high coll. rate)



Sampling zero-sound propagation

Can a BL theory develop spontaneously isosc. fluct. of correct amplitude?

- δI in homogeneous nucl. matter at low $T \Rightarrow$ 0-sound collective modes
- if unstable conditions \Rightarrow **amplification** \Rightarrow catastrophic process
- if T increases in time \Rightarrow 0-to-1st-sound [LARIONOV *et al* PRC61 (2000), KOLOMIETZ *et al* (1996)]
 \Rightarrow early times and small temperature restriction

- *linear-response* approx. \rightarrow assuming **small deviations** from f^0
 \Rightarrow *dispersion relation* from *linearised Vlasov* (no residual terms):

$$1 = \frac{g}{h^3} \frac{\partial U_k}{\partial \rho} \int \frac{\partial f^0}{\partial \epsilon} \frac{\mathbf{k} \cdot \mathbf{p}/m}{\omega_k + \mathbf{k} \cdot \mathbf{p}/m} d\mathbf{p}$$

at $T = 0$ integration restricted to the Fermi surface

self-consistency

$$\omega_k f_k + \mathbf{k} \cdot \frac{\mathbf{p}}{m} f_k - \frac{\partial f^0}{\partial \epsilon} \frac{\partial U_k}{\partial \rho} \mathbf{k} \cdot \frac{\mathbf{p}}{m} \rho_k = 0$$

at $T = 0$ eigenmodes f_k depend on states near the Fermi level ϵ_F

dispersion relation



$$-\frac{1}{F_0} = 1 - \frac{s}{2} \ln \left(\frac{s+1}{s-1} \right), \quad \text{sound velocity } s = \frac{\omega_k}{k v_F}$$

Landau parameter $F_0(\mathbf{k}) = \frac{3}{2} \frac{\rho^0}{\epsilon_F} \frac{\partial U_k}{\partial \rho}$ Fermi velocity

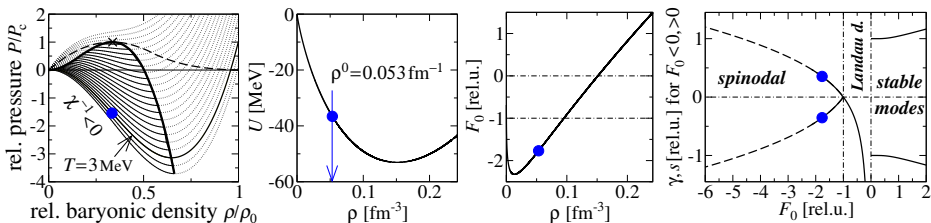
[LANDAU SovPhJETP 5 (1957)

KHALATNIKOV, ABRIKOSOV SovPhJETP 6 (1958)

COLONNA, CHOMAZ, RANDRUP

NPA567 (1994); PhREP 389 (2004)]

Instabilities in zero-sound conditions



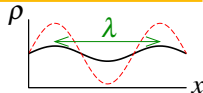
conditions of mechanical instability

$$\text{negative incompressibility } \chi^{-1} \equiv \rho \frac{\partial P}{\partial \rho} = \frac{2}{3} \rho \epsilon_F [1 + F_0(k=0)] < 0 \Rightarrow F_0(k=0) < -1$$

[POMARANCHUK SOVPHJETP 8 (1959)]

$$\Rightarrow \text{imaginary solutions } \gamma = is \text{ from } 1 + \frac{1}{F_0(k)} = \gamma \arctan \frac{1}{\gamma} \rightarrow |\gamma| = \frac{|\omega_k|}{k v_F} = \frac{1}{\tau_k k v_F}$$

\Rightarrow disturbances of wave number k get amplified with a growth time τ_k and a growth rate $\Gamma_k = 1/\tau_k$



Response intensity at zero-sound conditions, analytic

ultraviolet cutoff

- small k : linear $\Gamma \propto k$ evolution (the more matter to be relocated, the longer it takes)
- large k : small λ excluded as a function of the interaction range
→ Gaussian smearing σ in \mathcal{R} space

$$U \rightarrow U \otimes g(k), \text{ with } g(k) = e^{-\frac{1}{2}(k\sigma)^2}$$

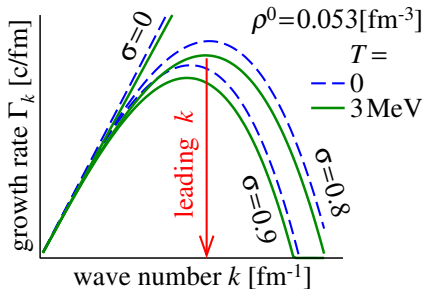
[COLONNA, CHOMAZ PRC49 (1994); KOLOMIETZ, SHLOMO PRC60 (1999)]

⇒ Γ spectra display a *leading k mode* (the fastest growing disturbance)

low, finite T

To fully explore the EOS (HIC ...) finite T should be included

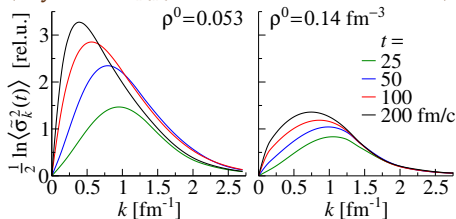
- Low- T expansion of chem.pot. $\eta(T > 0) \rightarrow \Gamma$ reduced



leading $k \Rightarrow$ most probable spinodal fragment size

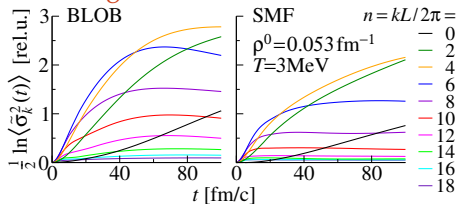
Response intensity at zero-sound conditions, BLOB

(asy-stiff, $\sigma_{NN} = \text{free} + \text{cut} 50 \text{mb}$, $l = 1 \text{fm}$)

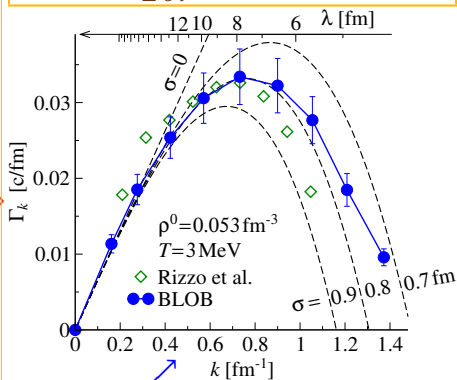


- $\rho^0 = 0.053 \text{ fm}^{-3} \rightarrow$ amplification \Rightarrow
- $\rho^0 = 1.4 \text{ fm}^{-3} \rightarrow$ Landau damping with significant amplitude

- evolution : earlier saturation in BLOB, large k combine into small k :



$$\Gamma_k = \frac{1}{2} \frac{\partial}{\partial t} \ln \langle \tilde{\sigma}_k^2(t) \rangle_{\text{paths}}$$



leading $\lambda \sim 8$ to 9 fm , $\tau_k \sim 200 \text{ fm/c}$

\Rightarrow Fragments arise with a size $A \approx \rho^0 \lambda^3$, in the region of neon, separating at around 200 fm/c

Isospin of emerging fragments in HIC

BLOB calculations in open systems :
 $N' - Z'$ distr. for forming clusters
 around C and Ne, before and during
 fragment formation,

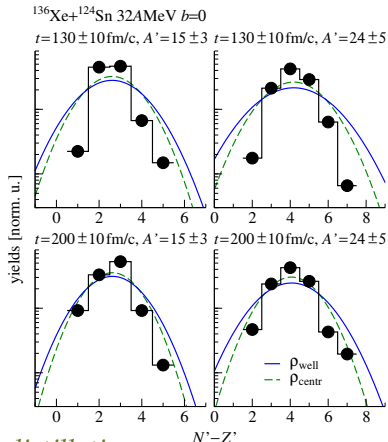
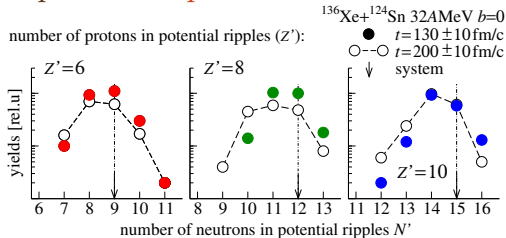
$$Y \approx \exp[-(\delta^2/A') C_{\text{sym}}(\rho)/T]$$

• underestimated but compensated by

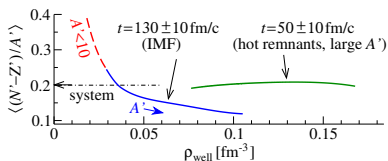
1) fluct. built out of equilibrium

→ higher coll.rate

2) particle evaporation



distillation :



Some BLOB highlights from spallation to HIC

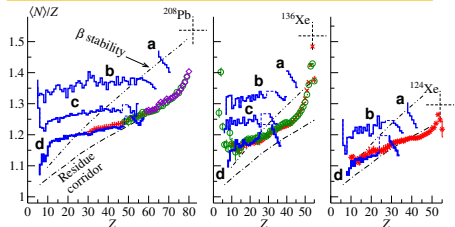
Fast binary splits from re-aggregation,

different from
asymm. fission

in $^{136}\text{Xe}+p$ 1A GeV :

[NAPOLITANI, COLONNA
PRC92 034607 (2015)]

Isospin "thermometry" in spallation :

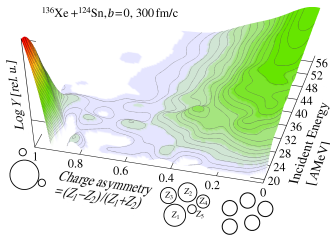


- BLOB, $^{208}\text{Pb}, ^{136}\text{Xe}, ^{124}\text{Xe}+p$ 1A GeV : (a) 200, (b) 400, (c) 700 fm/c, (d) cold
- + $^{208}\text{Pb}+^{124}\text{Xe}$ 1A GeV, *Enqvist 2001* × $^{136}\text{Xe}+^{124}\text{Xe}$ 1A GeV, *Henzlova 2008*
- $^{208}\text{Pb}+p$ 1A GeV, *Enqvist 2001* ○ $^{136}\text{Xe}+p$ 1A GeV, *P.N. 2007*
- ◇ $^{208}\text{Pb}+d$ 1A GeV, *Enqvist 2001* * $^{124}\text{Xe}+^{124}\text{Xe}$ 1A GeV, *Henzlova 2008*

[NAPOLITANI, COLONNA PRC92 034607 (2015)]

Bimodality in central HIC at Fermi E

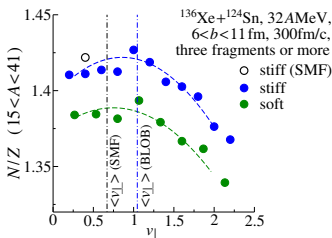
[NAPOLITANI,
COLONNA
PLB726 382
(2013)]



Neck formation and isospin migration

$$D_{iv}^{\rho} \nabla \rho \propto \frac{\partial S(\rho)}{\partial \rho} \nabla \rho$$

[COLONNA,
P.N., BARAN
NUC.CLUSTER
CORREL... (2017),



[BARAN2005; COLONNA2006; DITORO2006; RIZZO2008; ...]

Conclusions

BLE :

- No a priori assumption on the degree of equilibration
- BLOB : one-body theory based on a **full solution of the BLE**

Application to NM :

- BLOB : correct connection between MF pot. and unstable/stable is/iv modes tested in NM

Application to Open systems :

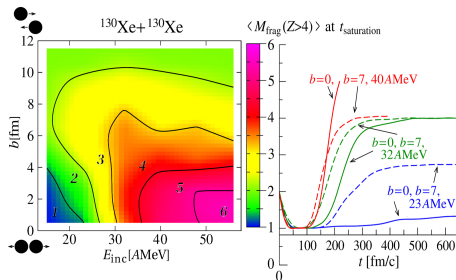
- tracking a variety of dynamical trajectories
- Fragment formation and recombination + isovector properties reasonably described over a large time interval

Dissipative HIC with BLOB

Fermi E :

large-ampl. phase-space fluct. +
in-medium dissipation + one-body
collective behaviour \Rightarrow

- fragments \rightarrow thresholds,
variety of mechanisms



- isospin currents as a function of t

Central / peripheral (BLOB) :

